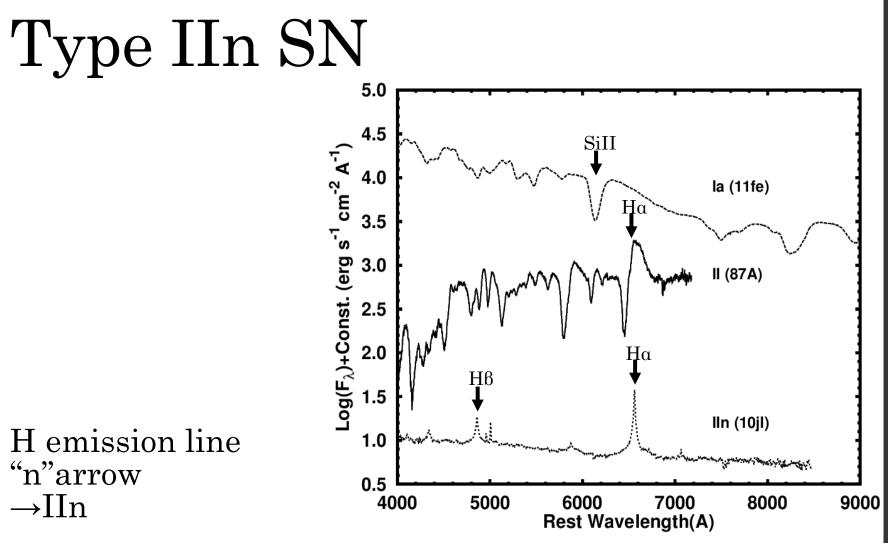
## 極めて大きな爆発エネルギーを 持つ超新星の観測的研究

大坪一輝、山中雅之、川端弘治、中岡竜也、川端美穂、 高木健吾、河原直貴、安部太晴(広島大学)秋田谷洋 (埼玉大学)



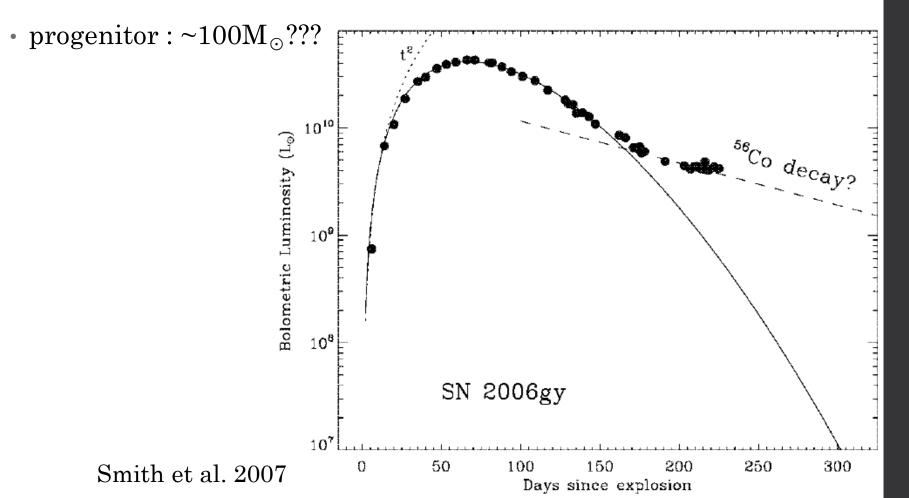


emission line

→circumstellar material (CSM) & ejecta interaction The origin of CSM is thought to be luminous blue variable (LBV) outburst

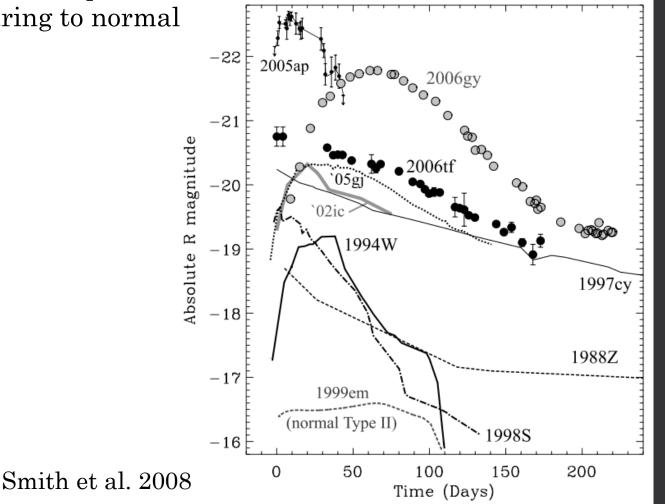
## SLSN-II

- Extremely luminous SN. (SN 2006gy ; -22mag)
- This can be explained by pair instability SN. (Smith et al. 2007 see also Yoshida-san's talk)

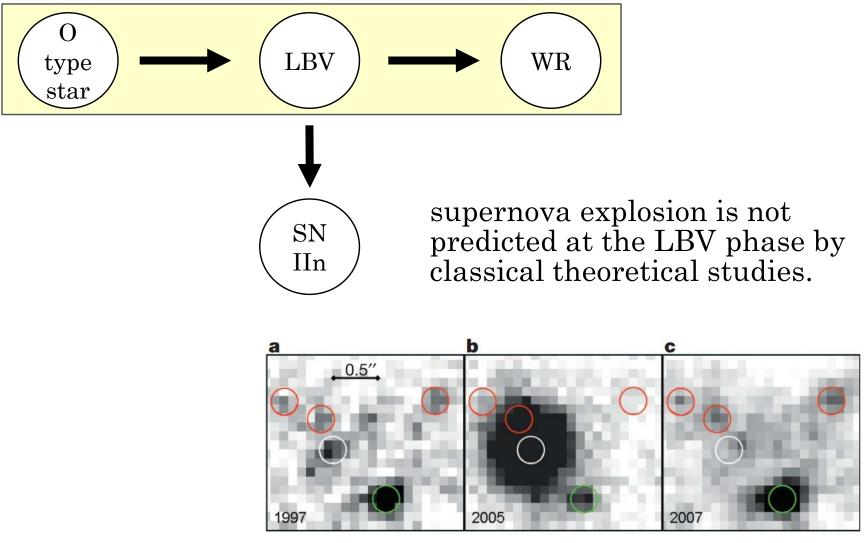


## SN IIn Light curves

- Their peak luminosities show the large diversity in order of 10,000.
- The light curves show quite slow evolution comparing to normal SNe II.



# Overview of the stellar evolution of massive stars



progenitor of SN 2005gl Gal-Yam et al. 2009

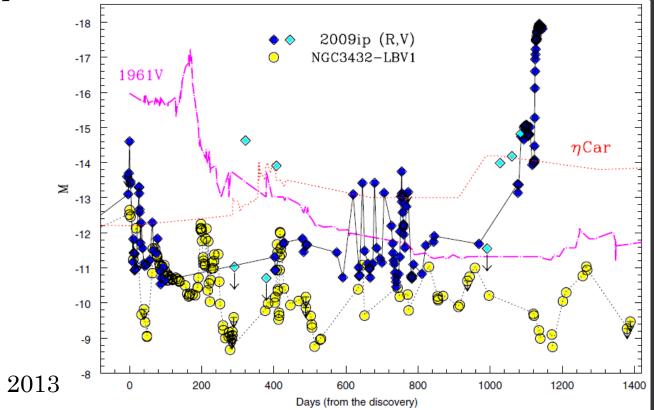
5

#### SN2009ip:rapid brightening after the LBV outburst

LBV outburst was discovered in 2009.

More luminous outburst was detected on October in 2012. Thereafter the rapid brightening was found and its peak reached to -18 mag.

From it, this eruption could be a SN IIn.



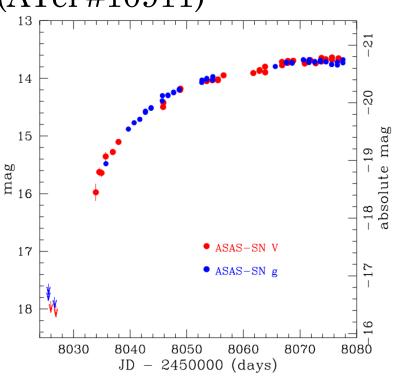
Pastorello et al. 2013

## SN 2017hcc

- Discovery:Oct-02-2017
- This SN was identified as a SN IIn.
- Polarization degree was reported to be 4.84%, which is quite large among SNe IIn. (ATel #10911)



Host galaxy : very faint (https://c1.staticflickr.com/5/4472/ 37224388064\_71d4839af0\_b.jpg)



Prieto et al. 2017

## Scientific goal

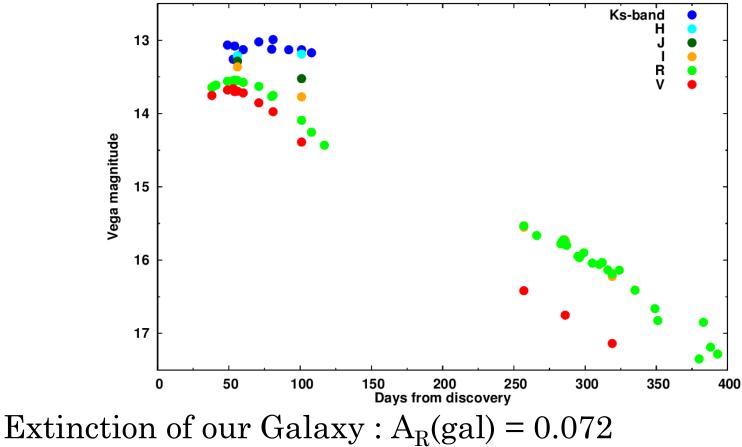
We discuss the unknown origin of the extremely luminous Type IIn SN with the Kanata telescope.

#### Observation and data reduction

- We carried out photometry (BVRIJHKs-band) and spectroscopy observation from November 09, 2017 to the last night, using the Kanata telescope (1.5m).
- Photometry was done with aperture.



## multi-band light curves



Extinction of host Galaxy : 0

R-band maximum is at 54 days after discovery. (13.5mag) Around 4 magnitude faded in 350 days.